

Diode Laser-Induced Fluorescence of Xenon Ion Velocity Distributions

Timothy B. Smith*, Bailo B. Ngom†, Jesse A. Linnell‡, and Alec D. Gallimore§

Plasmadynamics and Electric Propulsion Laboratory, University of Michigan, Ann Arbor, MI 48105

We present preliminary results from optogalvanic and laser-induced fluorescence (LIF) measurements of the nuclear-spin and isotopic structure of the $5d\ ^2F_{7/2} \rightarrow 6p\ ^2D_{5/2}^0$ transition of singly-ionized xenon (Xe II) at 834.7 nm. Ion velocity distributions deconvolved from $5d^4D_{7/2} \rightarrow 6p^4P_{5/2}^0$ optogalvanic spectra at 605.1 nm provided the Doppler broadening function for a computational model of the 834.7 nm absorption line. Ensemble averaging of optogalvanic spectra taken at 834.7 nm provided an experimental data set for estimation of the unknown isotopic shifts and lower-state ($5d\ ^2F_{7/2}$) nuclear-spin constants via a conjugate-gradient optimization routine. Deconvolution of 834.7 nm LIF spectra with the newly-modeled 834.7 nm cold-plasma spectrum returned an unusually broad velocity distribution, rather than the expected kinematic cooling effect. Comparison of a deconvolved ion energy distribution taken in a 500 V ion beam with retarding potential analyzer data showed bulk ion energy agreement within 1%, but only qualified similarity between the distributions. Grid misalignment during the LIF experiment may be responsible for the usually broad Doppler broadening. As a result, we are presently unable to reliably validate our deconvolution method at 834.7 nm.

I. Introduction

Laser-induced fluorescence (LIF) velocimetry of thruster plumes and internal flows requires a tunable, narrow-linewidth laser. Dye lasers, though capable of reaching a wide range of wavelengths, are complex, expensive and difficult to operate. However, dye lasers provide the only light source capable of singly-ionized xenon (Xe II) LIF driven by the $5d^4D_{7/2} \rightarrow 6p^4P_{5/2}^0$ absorption at 605.1 nm. This absorption is the only Xe II transition that currently has a well-established hyperfine structure (HFS), permitting accurate computational modeling of the LIF spectrum. As a result, deconvolution of the LIF at 605.1 nm with a simulated cold-plasma spectrum¹ is the only optical method currently available for high-spatial-resolution estimation of the Xe II velocity distribution.

External cavity diode lasers (ECDLs) are considerably less expensive and troublesome to operate than dye lasers, though more limited in wavelength range. Two neutral xenon (Xe I) absorptions ($6s[3/2]_2^0 \rightarrow 6p[3/2]_2$ at 823.2 nm and $6s[3/2]_1^0 \rightarrow 6p[1/2]_0$ at 828.0 nm) and one xenon ion (Xe II) absorption ($5d\ ^2F_{7/2} \rightarrow 6p\ ^2D_{5/2}^0$ at 834.7 nm) are readily probed by commercially-available diode lasers.² The 834.7 nm transition has proven to be especially popular for bulk velocity measurements in internal^{3–7} and external^{8,9} Hall thruster flows. Unfortunately, the isotopic shifts and lower state ($5d\ ^2F_{7/2}$) nuclear-spin structure constants are unknown for the 834.7 nm Xe II transition. Though the bulk velocity can be readily measured from the Doppler shift of the transition, the incomplete HFS model limits its usefulness in either deconvolving the ion velocity distribution or estimating its higher moments (such as the ion temperature) from the LIF spectrum.

In a series of Xe II optogalvanic spectroscopy (OGS) experiments carried out with a Hamamatsu L2783 xenon-neon galvatron, we have determined reasonable estimates of the missing lower-state ($5d\ ^2F_{7/2}$) nuclear-spin constants and isotopic shifts for the Xe II transition at 834.7 nm. Deconvolution of 49 OGS spectra

*Research Fellow, Aerospace Engineering, Senior Member AIAA.

†Graduate Student Research Assistant, Aerospace Engineering, Student Member AIAA.

‡FXB Fellow, Aerospace Engineering, Student Member AIAA.

§Professor, Aerospace Engineering, Associate Fellow AIAA

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recorded at 605.1 nm provided the Doppler broadening function for computational modeling of the 834.7 nm absorption line. A conjugate-gradient optimization routine then found values of the unknown HFS constants that minimized the mismatch between the model and the ensemble average of 61 OGS spectra at 834.7 nm.

In a concurrent series of 834.7 nm LIF measurements in the Cathode Test Facility (CTF) at the University of Michigan Plasmadynamics and Electric Propulsion Laboratory (PEPL), we attempted to demonstrate kinematic cooling of Xe II by electrostatic acceleration through the grids of a Commonwealth Scientific 3 cm ion source. Though the hoped-for kinematic cooling failed to appear, we were able to deconvolve the resulting LIF spectra with the newly-modeled cold-plasma spectrum at 834.7 nm. We compare the results of ion energy distributions estimated from LIF spectra and retarding potential analyzer (RPA) probing of the 3 cm ion source plume at a beam voltage of 500 V.

II. Theory

In LIF, the wavelength of a narrow-linewidth tunable laser is swept through an absorption line of a plasma species, such as singly-ionized xenon (Xe II). Focusing optics direct this laser beam along a “beamwise” direction vector $\hat{\mathbf{k}} \equiv \mathbf{k}/|\mathbf{k}|$, where \mathbf{k} is the beam’s wave vector. Collection optics on another (usually perpendicular) axis sample the fluorescence emitted by absorbing particles in the interrogation volume. This interrogation volume, an ellipsoid defined by the intersecting beams, can be less than a millimeter on a side. Wavelength filtering and lock-in detection keep the LIF signal from being swamped by the bright light background typical of EP devices.

A. Line broadening

Consider a light source with vacuum wavelength λ and frequency $\nu = c/\lambda$. An observer who is stationary with respect to the light source will see light at the same frequency ν . An observer moving towards the light source will see a bluer (*i.e.*, higher-frequency) light than the stationary viewer, while an observer moving away from the light source will see a redder (*i.e.*, lower-frequency) light.

In LIF, this Doppler effect appears as a shift in the resonant frequency ν_0 as the laser is scanned over a very short frequency range. The change in photon frequency $\Delta\nu = \nu - \nu_0$ for a particle with velocity \mathbf{v} passing through a light beam of wave vector \mathbf{k} is

$$\Delta\nu = -\frac{\mathbf{k} \cdot \mathbf{v}}{2\pi}. \quad (1)$$

Given a beamwise velocity component $v_k \equiv \mathbf{v} \cdot \hat{\mathbf{k}}$, Eqn. 1 gives the beamwise velocity-to-frequency transformations

$$\nu = \left(1 - \frac{v_k}{c}\right) \nu_0 \quad (2)$$

$$v_k = \left(1 - \frac{\nu}{\nu_0}\right) c. \quad (3)$$

A swarm of particles with a normalized velocity distribution $f(\mathbf{v})$ will also “see” the frequency of incoming photons shifted by the relative velocity of the particle in the direction of the photon. The resulting Doppler lineshape will be shifted by the beamwise bulk velocity $u_k = \langle \mathbf{v} \cdot \hat{\mathbf{k}} \rangle$ and broadened by the thermal width of the beamwise distribution $f(v_k)$. The generalized Doppler lineshape $d(\nu)$, when properly normalized so that

$$\int_{-\infty}^{\infty} d(\nu) d\nu = 1, \quad (4)$$

is given by

$$d(\nu) = \frac{c}{\nu_0} f\left(\left[1 - \frac{\nu}{\nu_0}\right] c\right). \quad (5)$$

When $f(v_k)$ is a one-dimensional stationary Maxwellian of the form

$$f_m(v_k) = \left(\frac{M}{2\pi kT}\right)^{1/2} \exp\left(-\frac{Mv_k^2}{2kT}\right), \quad (6)$$

Eqn. 5 takes the familiar form¹⁰

$$d_m(\nu) = \frac{c}{\nu_0} \left(\frac{M}{2\pi kT} \right)^{1/2} \exp \left(-\frac{Mc^2}{2kT} \left[\frac{\nu - \nu_0}{\nu_0} \right]^2 \right). \quad (7)$$

In species with no hyperfine structure (HFS), the LIF spectrum $i(\nu)$ is the convolution of the Doppler broadening with the natural (or lifetime) broadening function

$$l(\nu) = \frac{\Delta\nu_n}{2\pi} \frac{1}{(\nu - \nu_0)^2 + (\Delta\nu_n/2)^2} \quad (8)$$

where $\Delta\nu_n = A_j/(2\pi)$ is the natural linewidth for state j . This Lorentzian function is usually much narrower for warm gases than the Doppler broadening. In this case, the natural broadening can be replaced by a Dirac delta function,

$$i(\nu) = d(\nu) \otimes l(\nu) \approx d(\nu) \otimes \delta(\nu) = d(\nu), \quad (9)$$

letting us directly transform the LIF spectrum into an excellent approximation of the beamwise velocity distribution. The similarity between LIF spectra and $f(v_k)$ is good enough that LIF measurements in barium^{11–13} and argon¹⁴ plasmas have been reported as velocity distributions.

B. Hyperfine structure

In xenon (and other species with HFS), the LIF spectrum can be modeled as the convolution of the hyperfine splitting function $h(\nu)$ with natural and Doppler broadening functions,

$$i(\nu) = h(\nu) \otimes l(\nu) \otimes d(\nu). \quad (10)$$

The hyperfine splitting $h(\nu)$ is a series of n Dirac delta functions,

$$h(\nu) = \sum_{j=1}^n p_j \delta(\nu - \nu_j). \quad (11)$$

where ν_j is the j^{th} hyperfine line center and p_j is the j^{th} hyperfine line intensity. The $5d^4D_{7/2} \rightarrow 6p^4P_{5/2}^0$ and $5d^2F_{7/2} \rightarrow 6p^2D_{5/2}^0$ transitions of Xe II has $n = 19$ such hyperfine splittings.

Hyperfine structure consists of two components. Isotopic splitting is caused by the mass and nuclear volume differences between isotopes in the sample. Nuclear-spin splitting is caused by the coupling between the nuclear spin \mathbf{I} and the total electronic angular momentum \mathbf{J} . For the sake of brevity, we will henceforth use the term ‘‘HFS constants’’ to denote the combined set of isotopic shifts and nuclear-spin interaction constants.

There are nine stable isotopes of xenon, seven of which have natural abundances greater than one percent. Each of these isotopes causes a slightly different transition energy E_{ij} , resulting in isotopic splitting. Both of the computational models in this study use naturally-occurring xenon isotopic abundances,¹⁵ while we take isotopic shifts (conventionally defined relative to the center frequency for ¹³²Xe) for the $5d^4D_{7/2} - 6p^4P_{5/2}^0$ transition at 605.1 nm from the fast ion-beam LIF surveys of Bingham *et al.*¹⁶ and Borghs *et al.*¹⁷

Seven of the nine stable isotopes of xenon have even atomic mass, resulting in no nuclear spin \mathbf{I} . The two isotopes with an odd atomic mass, however, have non-zero nuclear spin quantum numbers I . The lighter isotope, ¹²⁹Xe, has $I = 1/2$, while ¹³¹Xe has $I = 3/2$. These non-zero nuclear spins cause nuclear-spin splitting of the atomic energy levels. This nuclear-spin splitting is 1 to 2 orders of magnitude broader than the isotopic splitting, and provides most of the characteristic shape of the $5d^4D_{7/2} - 6p^4P_{5/2}^0$ line at 605.1 nm. The $5d^2F_{7/2} \rightarrow 6p^2D_{5/2}^0$ line at 834.7 nm is considerably narrower, but nuclear-spin splitting also has a disproportionate influence on the lineshape.

The total angular momentum $\mathbf{F} = \mathbf{I} + \mathbf{J}$ takes quantum number values

$$F = I + J, I + J - 1, \dots, |I - J| \quad (12)$$

where J is the total electronic angular momentum quantum number.¹⁸ The extra term energy due to nuclear-spin splitting is given by¹⁹

$$E_{\text{NSS}} = A \frac{C}{2} + BD, \quad (13)$$

where A is the nuclear magnetic dipole interaction constant, B is the nuclear electric quadrupole interaction constant, and the terms

$$C = F(F + 1) - I(I + 1) - J(J + 1) \quad (14)$$

and

$$D = \frac{(3C/4)(C + 1) - I(I + 1)J(J + 1)}{2I(2I - 1)J(2J - 1)} \quad (15)$$

contain the nuclear spin-orbit interactions. Table 1 presents Bröstrom's nuclear-spin structure constants for the Xe II $5d^4D_{7/2}$, $6p^4P_{5/2}^0$ and $6p^2D_{5/2}^0$ energy levels.²⁰

Table 1. Published nuclear-spin structure constants for the Xe II transitions at 605.1 nm and 834.7 nm.

Level	A_{129} (MHz)	A_{131} (MHz)	B_{131} (MHz)
$5d^4D_{7/2}$	-502 ± 4	150 ± 2	70 ± 13
$6p^4P_{5/2}^0$	-1633 ± 11	487 ± 4	-129 ± 16
$6p^2D_{5/2}^0$	-1387 ± 9	409 ± 2	-117 ± 11

The transition rule for nuclear-spin splitting is $\Delta F \equiv F - F' = [0, \pm 1]$, where F is the upper and F' is the lower state's total angular momentum quantum number. (The zero-zero transition is forbidden, $F = 0 \not\rightarrow F' = 0$.) Since the total electronic angular momentum quantum numbers are the same for both transitions ($J = 7/2 \rightarrow J' = 5/2$), both lines have the same range of allowable total angular momentum quantum numbers. As a result, Figure 1 shows the fine structure and nuclear-spin splitting for both the $5d^4D_{7/2} - 6p^4P_{5/2}^0$ and the $5d^2D_{5/2}^0 - 6s^2P_{3/2}$ lines.

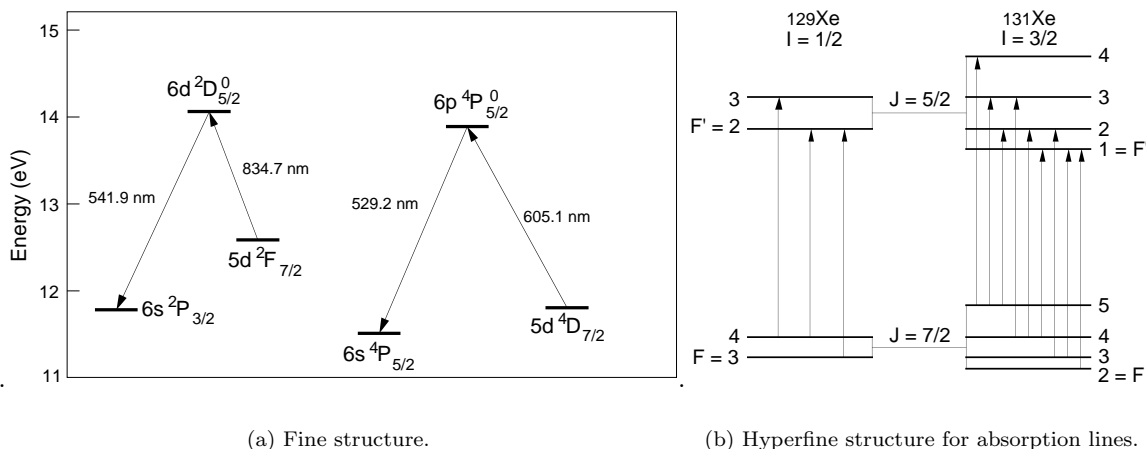


Figure 1. Fine and hyperfine structure for 605.1 nm and 834.7 nm Xe II transitions.

The relative intensity of each nuclear-spin split component is given for a $J \rightarrow J - 1$ transition by²¹

$$I(F \rightarrow F - 1) \propto \frac{P(F)P(F - 1)}{F} \quad (16)$$

$$I(F \rightarrow F) \propto \frac{(2F + 1)}{F(F + 1)} P(F)Q(F) \quad (17)$$

$$I(F - 1 \rightarrow F) \propto \frac{Q(F)Q(F - 1)}{F} \quad (18)$$

where $P(F) = (F + J)(F + J + 1) - I(I + 1)$ and $Q(F) = I(I + 1) - (F - J)(F - J + 1)$.

For the isotopes with even mass numbers (i.e., without nuclear-spin splitting), the line intensity p_j is linearly proportional to the naturally-occurring abundance for each isotope. The line intensities of ^{129}Xe

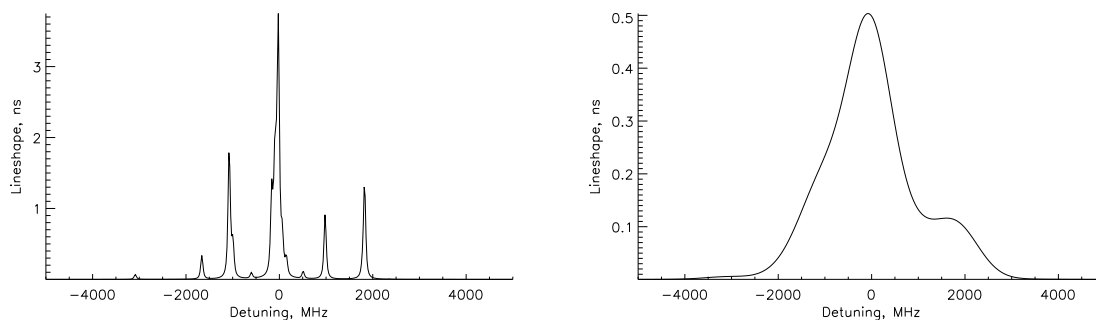
and ^{131}Xe are linearly proportional to the product of the isotopic abundance and the relative intensity of the nuclear-spin split components. It is worth noting that the relative intensity is *not* a function of the HFS constants, but is the same for both the $5d\ ^4D_{7/2} - 6p\ ^4P_{5/2}^0$ and the $5d\ ^2F_{7/2} \rightarrow 6p\ ^2D_{5/2}^0$ lines. This fact considerably simplifies the task of estimating HFS constants from the 834.7 nm data.

Table 2 presents naturally-occurring xenon isotopic abundances and shifts for the $5d\ ^4D_{7/2} - 6p\ ^4P_{5/2}^0$ transition taken from the fast ion-beam LIF surveys of Bingham *et al.*¹⁶ and Borghs *et al.*¹⁷

Table 2. Isotopic abundance and shift relative to ^{132}Xe for the $5d\ ^4D_{7/2} - 6p\ ^4P_{5/2}^0$ transition at 605.1 nm.

Mass (amu)	124	126	128	129	130	131	132	134	136
Abundance (%)	0.0096	0.009	1.92	26.4	4.1	21.1	26.9	10.4	8.9
Shift (MHz)	336.6	252.4	172.0	113.7	83.6	16.7	0.0	-75.8	-140.9

C. Deconvolution



(a) Cold-plasma spectrum $c(\nu)$.

(b) Warm-plasma spectrum $w(\nu)$, $T = 1255$ K.

Figure 2. Cold-plasma and warm-plasma spectra for the Xe II $5d^4D_{7/2} - 6p^4P_{5/2}^0$ line at 605.1 nm.

Figure 2(a) shows the 605.1 nm absorption spectrum for a perfectly cold stationary plasma, where the velocity distribution $f(\mathbf{v}) = \delta(\mathbf{v})$. This can be described by the convolution

$$c(\nu) = h(\nu) \otimes l(\nu), \quad (19)$$

so that Eqn. 10 becomes (in the absence of noise)

$$w(\nu) = c(\nu) \otimes d(\nu). \quad (20)$$

Given the Fourier transforms $C(\tau)$ and $D(\tau)$ of the cold-plasma and Doppler broadening functions $c(\nu)$ and $d(\nu)$, the convolution theorem states that Eqn. 20 is equivalent to the product

$$I(\tau) = C(\tau)D(\tau) \quad (21)$$

where $I(\tau)$ is the Fourier transform of $i(\nu)$. It then follows that the Doppler broadening transform estimate

$$\hat{D}_s(\tau) = I(\tau)/C(\tau) \quad (22)$$

is equivalent to the deconvolution

$$\hat{d}_s(\nu) = i(\nu) \oslash c(\nu) \quad (23)$$

which can be converted to a beamwise velocity distribution estimate by the transformation

$$\hat{f}(v_k) = \frac{1}{\lambda_0} \hat{d}_s \left(\left[1 - \frac{v_k}{c} \right] \nu_0 \right). \quad (24)$$

This simple inverse filter approach, though theoretically sound, quickly collapses for realistic signal-to-noise ratios. Deconvolution methods tend to preferentially amplify high-frequency noise.²² Low-pass filtering of the deconvolved velocity distribution estimate $\hat{f}(v_k)$ with a Gaussian function of bandwidth \mathcal{T} yields an improved Doppler broadening transform estimate

$$\hat{D}_g(\tau) = \frac{\exp(-[\tau/\mathcal{T}]^2)}{C(\tau)} I(\tau) \quad (25)$$

that is equivalent to convolving the simple Doppler broadening function estimate with a scale-similar Gaussian,

$$\hat{d}_g(\nu) = \pi\mathcal{T}^2 \exp(-[\pi\mathcal{T}\nu]^2) \otimes \hat{d}_s(\nu). \quad (26)$$

We have found²³ that a bandwidth of $2 \text{ ns} \leq \mathcal{T} \leq 3 \text{ ns}$ reduces the noise amplification factor to an acceptable level for LIF and OGS spectra with reasonable signal-to-noise ratios ($S/N \geq 30$), at the cost of discarding some high-frequency information along with the noise.

D. Conjugate-gradient method

Deconvolution of the velocity distribution from LIF and OGS spectra requires a solid model of the cold-plasma spectrum $c(\nu)$ for the absorption line. In the absence of published isotopic shifts and lower-state nuclear-spin constants for the $5d \ ^2F_{7/2} \rightarrow 6p \ ^2D_{5/2}^0$ transition at 834.7 nm, we used the conjugate gradient (CG) method²⁵ to recover the HFS constants from OGS spectra.

The CG method is an iterative procedure that recovers a local extremum of the function of interest $e(\vec{x})$ by successive searches along steepest descent directions in the parameter space denoted by \vec{x} . Consider the absorption line model $w(\vec{x}, \nu_j)$, where the parameter space

$$\vec{x} = [A_{129}, A_{131}, B_{131}, \Delta\nu_{128}, \Delta\nu_{129}, \Delta\nu_{130}, \Delta\nu_{131}, \Delta\nu_{134}, \Delta\nu_{136}] \quad (27)$$

includes the two lower-state nuclear magnetic dipole interaction constants (A_{129} and A_{131}), the one lower-state nuclear electric quadrupole interaction constant (B_{131}), and six isotopic shifts ($\Delta\nu_{128}$, $\Delta\nu_{129}$, $\Delta\nu_{130}$, $\Delta\nu_{131}$, $\Delta\nu_{134}$, and $\Delta\nu_{136}$). In the optimization problem of searching for the 834.7 nm HFS constants contained in \vec{x} , we want to minimize the error function

$$e(\vec{x}) = \sum_{j=1}^n [i(\nu_j) - w(\vec{x}, \nu_j)]^2 \quad (28)$$

where the measured OGS spectrum $i(\nu)$ consists of n measurements of the OG cell current $i(\nu_j)$ at n frequency shifts ν_j .

The first step of the procedure is a guess of the solution, \vec{x}_0 . This initial guess must be reasonably close to the target solution, or else the method may return a solution that corresponds to a local minimum, rather than the desired global minimum. We found that a good initial guess could be found from the intensity-weighted nuclear-spin splitting magnitudes

$$\theta_{129}^2 = \frac{1}{3} \sum_{i=1}^3 (p_i \Delta\nu_i)^2 \quad (29)$$

$$\text{and } \theta_{131}^2 = \frac{1}{9} \sum_{i=1}^9 (p_i \Delta\nu_i)^2 \quad (30)$$

where p_i is the relative intensity of the hyperfine line and $\Delta\nu_i = (E'_{\text{nss}} - E_{\text{nss}})/h$ is the hyperfine line shift. Since the 834.7 nm line shows much less hyperfine broadening than the 605.1 nm line, we used HFS constants that minimize $\theta_{129}(A_{129})$ and $\theta_{131}^2(A_{131}, B_{131})$ to specify \vec{x}_0 .

The initial steepest-descent vector \vec{d}_0 from \vec{x}_0 is the negative of the gradient of the error function at \vec{x}_0 ,

$$\vec{d}_0 = -\nabla e(\vec{x}_0). \quad (31)$$

Note that \vec{d}_0 is not necessarily a vector pointing towards the desired minimum. We can only expect that there exists some local minimum $\vec{x}_{m,0}$ along the intersection defined by a hyper-plane \mathbf{P}_0 containing \vec{d}_0 and

the hyper-surface $e(\vec{x})$ defining our error function. The minimum along this first hyperplane, which becomes the second guess at the solution \vec{x}_1 , is given by

$$\vec{x}_{m,0} = \vec{x}_0 + \alpha_0 \vec{d}_0 = \vec{x}_1 \quad (32)$$

where the coefficient α_0 is an optimal displacement length found by the requirement that the new steepest-descent vector \vec{d}_1 must be orthogonal to the initial gradient,

$$-\nabla e(\vec{x}_1) \cdot \vec{d}_0 = 0. \quad (33)$$

Substituting Eq. 32 into Eq. 33 yields the condition

$$-\nabla e(\vec{x}_0 + \alpha \vec{d}_0) \cdot \vec{d}_0 = 0. \quad (34)$$

This, and the optimal displacement lengths corresponding to subsequent steepest-descent vectors, can be iteratively solved via the secant method,²⁵

$$\alpha_i = -\sigma_i \frac{\nabla e(\vec{x}_i) \cdot \vec{d}_i}{\nabla e(\vec{x}_i + \sigma_i) \cdot \vec{d}_i - \nabla e(\vec{x}_i) \cdot \vec{d}_i} \quad (35)$$

$$\text{and } \sigma_i = -\alpha_i. \quad (36)$$

Convergence on a value of α_i by Eq. 36 prompts the calculation of a new steepest-descent vector. In order to avoid retracing prior search paths, we use the Fletcher-Reeves formula²⁴

$$\vec{d}_{i+1} = -\nabla e(\vec{x}_{i+1} + \beta_{i+1} \vec{d}_i) \quad (37)$$

where the constant

$$\beta_{i+1} = \frac{\nabla e(\vec{x}_{i+1}) \cdot \nabla e(\vec{x}_i)}{\nabla e(\vec{x}_i) \cdot \nabla e(\vec{x}_i)} \quad (38)$$

for $i > 0$ and $\beta_0 = 0$.

The primary stopping condition for our CG code is given when the square of the change in \vec{x} drops below a fixed magnitude ϵ ,

$$|\vec{x}_{i+1} - \vec{x}_i|^2 > \epsilon. \quad (39)$$

Given that $\vec{x}_{i+1} = \vec{x}_i + \alpha_i \vec{d}_i$, the stopping condition for convergence becomes

$$\alpha_i^2 \vec{d}_i \cdot \vec{d}_i < \epsilon, \quad (40)$$

which requires no *a priori* knowledge of the next parameter vector \vec{x}_{i+1} .

E. Optogalvanic spectroscopy

Optogalvanic spectroscopy records changes in the current of a constant-voltage discharge as a laser beam passing through the plasma is tuned through an absorption line of one of the constituent plasma species. Since the current in an unmagnetized discharge is dominated by electron flow, the source of the optogalvanic effect is increased electron density due to electron-impact ionization. For singly-ionized xenon, the relevant reaction is



Since upper states of Xe II have term energies that are closer to the second ionization potential than lower states, they tend to have a higher probability of electron-impact ionization. Thus, absorption increases the upper-state population and depletes the lower-state population; the net change in the electron density is positive, and shows up as an AC component of the galvatron current at the laser chopping frequency.

This AC galvatron current can be modeled as the sum of signal, noise, and background components:

$$I_g(\nu) = I_{og}(\nu, P_L) + n(\nu) + I_p(P_L). \quad (42)$$

The optogalvanic current $I_{og}(\nu, P_L)$, the signal of interest, is a function of the laser frequency ν and the laser power at the galvatron input window P_L . Following the semi-empirical method we developed to model the saturated absorption lineshape $i_s(\nu, P_L)$,²³ the optogalvanic current can be modeled by

$$I_{og}(\nu, P_L) \propto \frac{\beta P_L i(\nu)}{1 + \beta P_L i(\nu)} \quad (43)$$

where $i(\nu)$ is the unsaturated absorption lineshape and the saturation coefficient β is a experimentally-determined constant for a given transition and set of plasma parameters.

Though the noise component $n(\nu)$ can be diminished by ensemble averaging of multiple scans, we have found that there is an irreducible background current $I_p(P_L)$ for any optogalvanic setup. This suggests that there is another source of increased discharge current at the chopping frequency. Good centering of the laser beam through the galvatron electrodes tends to increase the signal-to-background ratio, as does reduction of the beam diameter via improved focus and/or a field stop.

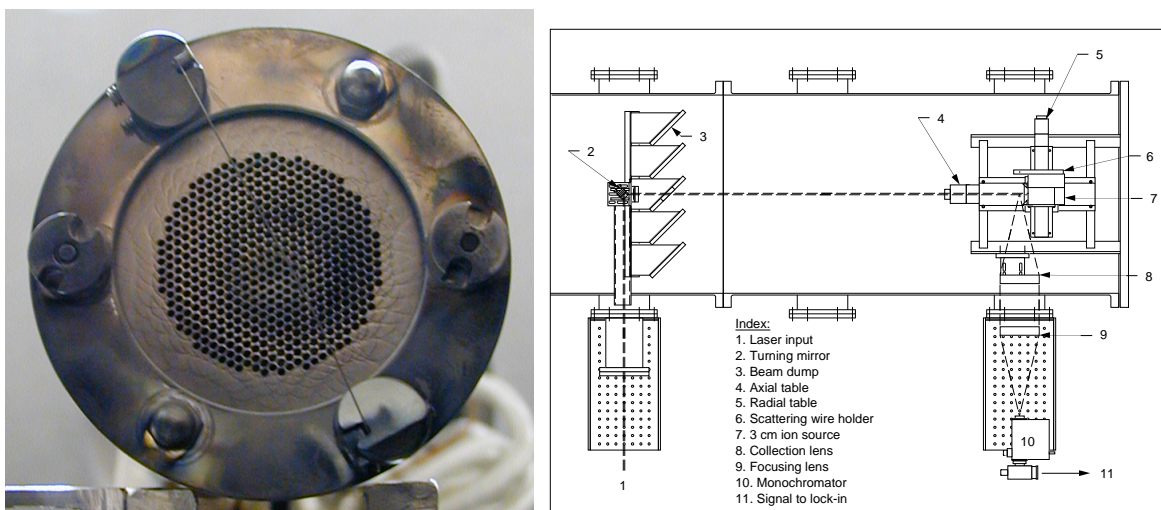
We believe that the background current $I_p(P_L)$ can be explained as photoelectric emission, induced by the absorption at the cathode surface of laser scattering and Fresnel reflections within the galvatron. This scales linearly with the photon flux to the cathode surface, $I_p(P_L) \propto P_L$; thus, for a given experimental setup, the maximum signal-to-background ratio becomes

$$\frac{I_{og}}{I_p} \propto \frac{\beta \max[i(\nu)]}{1 + \beta P_L \max[i(\nu)]} \quad (44)$$

which replicates the observed decrease of the signal-to-background ratio with laser power.

III. Experimental apparatus

A. Facility



(a) Photograph of the 3 cm ion source.

(b) Beam-handling and collection optics, plan view.

Figure 3. Experimental setup in the Cathode Test Facility.

All LIF experiments were carried out in the Cathode Test Facility (CTF) at PEPL. The CTF is a $\phi 0.6 \text{ m} \times 2 \text{ m}$ aluminum-walled cylindrical tank. Ten $\phi 20 \text{ cm}$ flanges allow considerable flexibility in the mounting of optical windows, pressure gauges, feedthroughs, and diagnostics. A five-panel graphite beam dump protects the cryopump and suppresses back sputtering caused by the ion beam.

During operation, an Edwards E2M-30 two-stage mechanical pump roughs the CTF at 135 cfm to approximately 60 mTorr. At this point, the mechanical pumping system is sealed and turned off, and a $\phi 50$

cm CVI TM500 cryopump with a measured xenon pumping speed of 1,500 l/s takes over. Base pressure for the facility is 2×10^{-8} Torr, as measured with a Varian model 421 cold cathode gauge.

The ion source was mounted on two crossed translation tables, a NEAT TMS-1100-SM single-axis shuttle with 11 inches of travel and a NEAT RMS-800-SM single-axis shuttle with 8 inches of travel. Both translation stages are PC-controlled by a custom LabView VI, with locational resolution on the order of 0.1 mm.

Figure 3(a) shows our Commonwealth Scientific $\phi 3$ cm gridded ion source, a commercial unit capable of operation at up to 1500 V. The outer body of the ion source is a grounded $\phi 63.5 \times 84.8$ mm shell of 1.6 mm thick stainless steel. The screen and accelerator grids are $\phi 43.5 \times 1$ mm discs with three mounting lugs and a laser-drilled hexagonal pattern of $\phi 1.0$ mm apertures with 1.32 between centers. Alumina spacers (when uncoated by sputtering products) isolate both grids and maintain the grid spacing. A Unit UFC-7301 mass flow controller regulates 0 to 10 sccm of Xe flow into the discharge. A Kikisui PAD 16-10L supply powers the $\phi 0.25 \times 60$ mm tungsten discharge cathode filament, while a Hewlett-Packard HP 6271 supply maintains the discharge current and voltage. Both of these supplies are floated atop a Glassman FC01.5P80 1500-V, 80-mA power supply, while a Sorenson DCS 33-33 supply powers the $\phi 0.25 \times 80$ mm neutralizer cathode filament.

B. Lasers and external optics

The primary laser system used in this experiment, a TUI Optics TA-100/830 tapered-amplifier diode laser, has not previously been used at PEPL. This system has a nominal linewidth of 2 MHz, a tuning range of 820 nm to 836 nm, and a mode-hop-free range of over 10 GHz. Though capable of output powers in excess of 500 mW, off-peak efficiency drops and Faraday isolation of the tapered amplifier reduce the available output power at the upper end of the system wavelength range (835 ± 1 nm) to approximately 250 mW.

The master oscillator for this system is an ECDL. Coarse wavelength control is provided by Peltier cooling of the ECDL and micrometer adjustment of the ECDL feedback grating angle. PC control of the fine tuning is provided by a LabView VI, which uses the auxiliary output of a Stanford SR810 lock-in amplifier to send a control voltage to a piezoelectric transducer (PZT) driving the ECDL feedback grating. Feedback circuitry in the ECDL power supply decreases the master oscillator diode current during a PZT sweep to maximize the mode-hop-free range. In order to include the effect of changing laser output power, we normalize our reported LIF and OGS spectra by a quadratic curve-fit to the master oscillator power. It is important to note that this normalization carries an implicit assumption that the absorption line is not saturated. Steering mirrors inside the TA-100 enclosure couple the ECDL output into the tapered amplifier diode, which boosts the ECDL output power by a factor of 10 while maintaining the master oscillator wavelength and linewidth.

Figure 4(a) shows a schematic of the diode laser table contents. A wedged beamsplitter sends approximately 5% of the TA-100 output beam to a steering mirror, and thence to a Burleigh WA-1000 wavemeter. This wavemeter permits wavelength measurement with a 1 pm resolution. Oversampling of the PZT control voltage permits greater wavelength resolution by a quadratic curve-fit to the wavemeter output, but does not improve the absolute bias of the laser wavelength measurement.

The second-surface reflection from the wedged beamsplitter is currently sent to a beam dump. In future experiments, we plan to direct this beam into a recently-refurbished Burleigh SApplus spectrum analyzer). This 2-GHz etalon, which has a finesse of 256, will provide repeatable frequency markers at 2-GHz intervals, providing reference markings by which individual 10-GHz scans can be “stitched” together into a reasonably seamless scan over many tens of GHz.

The remaining 90% of the beam passes through a Stanford SR541 optical chopper, permitting phase-lock amplification of subsequent OGS and LIF signals.

The secondary system, a Coherent 899-29 Autoscan ring dye laser, has been used for LIF measurements for the past six years at PEPL. This PC-controlled system has a nominal linewidth of 500 kHz, tuning repeatability of 50 MHz and a scanning range of over 100 GHz (in 10 GHz segments). PC-controlled scanning and data collection are synchronized by the Autoscan software. Pumping for this dye laser is provided by a Coherent SBRC-R-DBW20/4 argon-ion laser, with a nominal broadband power rating of 25 W. With the intercavity assembly (ICA) removed, the dye laser can generate up to 2 W of tunable broadband light using Rhodamine-6G dye at 605 nm. With the ICA installed, the same system can provide anywhere from 300 to 450 mW of narrow-linewidth light at 605.1 nm.

In order to combine beams from both laser systems on a single beamline, we designed the beam-mixer shown in Figure 4(b). The heart of this is a polarizing beamsplitter. Run in the usual orientation, it takes a randomly-polarized input beam and outputs two orthogonally-directed beams, each polarized at right angles to one another. In our setup, the beamsplitter takes two orthogonally-polarized input beams and combines

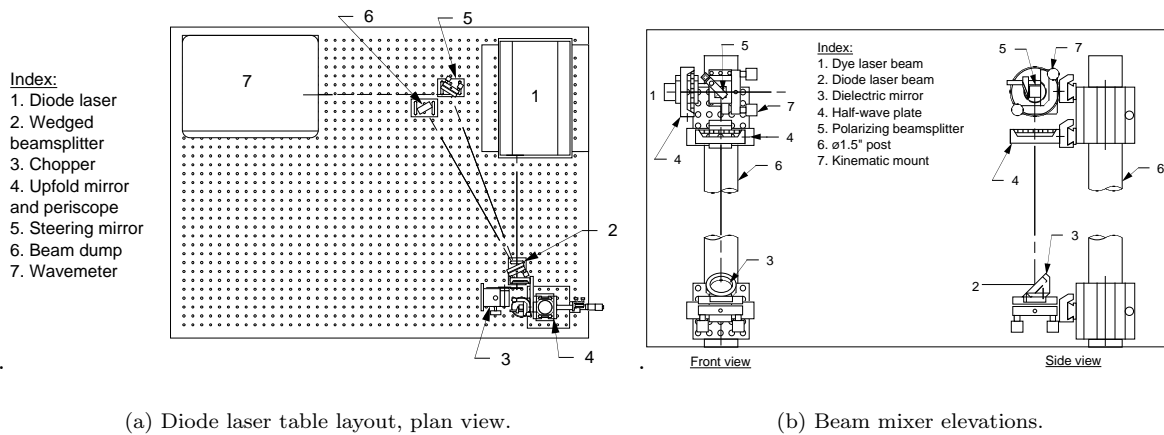


Figure 4. Laser beam mixing, modulation and sampling.

them on a single beamline. Half-wave plates, which are mounted in rotation stages between the lasers and the beamsplitter, ensure that the polarization direction of the incoming beams is properly aligned for maximum transmission through the beamsplitter. A kinematic mount permits fine control of the diode laser beam direction, ensuring that the two beams remain overlapped throughout the experimental setup. A $\phi 25$ mm, $f = 2$ m lens placed at the beam-mixer output places the combined beam focal point within a few cm of the interrogation volume within the CTF.

Since the dye laser system was not working well enough for reliable LIF during this experiment, we substituted a simple helium-neon laser. Though this is not useful for LIF, it nevertheless provided a valuable tool for visible-light alignment of the laser input and collection optics axes.

Figure 3(b) shows the beam-handling and collection optics setup used for LIF experiments in the CTF. A square 25 mm mirror on a kinematic mount, attached within an enclosure tube behind the beam dump, redirects the combined laser beams through a $\phi 25$ mm protective window toward the ion source grid. A $\phi 1 \times 80$ mm section of steel music wire, mounted vertically in a tube placed to one side of the ion source, facilitates alignment of the collection optics. After moving the alignment wire into the laser beam path, we adjust the $\phi 100$ mm, $f/2.5$ collection lens, sending a collimated beam of scattered light through the CTF window. During experiments, the collimated fluorescence follows the same path.

The collimated fluorescence from the ion source plume is focused by a second $\phi 100$, $f/2.5$ lens onto the entrance slits of a Spex H-10 monochromator with a Hamamatsu 928 PMT. This monochromator acts as a linewidth filter centered on the 541.9 nm fluorescence line. By holding a second-surface mirror flat against the monochromator entrance slits, we can use retroreflection of scattered light from the alignment wire to ensure proper alignment of the monochromator with the collection optics axis. The PMT signal is passed to a Stanford SR810 lock-in amplifier, which isolates the fluorescence component of the optical signal entering the PMT.

C. Optogalvanic cell

Figure 5 shows the optogalvanic cell setup used on the CTF. A wedged beamsplitter, located just before the CTF input window, sends approximately 4% of the combined beams to a steering mirror, and thence through the optogalvanic cell. A filter holder placed in the beam path permits neutral density filter attenuation of the incoming laser beam power for saturation studies.

The optogalvanic cell used in this study is a Hamamatsu L2783-42 XeNe-Mo galvatron. The core of the galvatron is a $\phi 6.25$ mm pair of cylindrical Mo electrodes, centered in a $\phi 25 \times 120$ mm glass cylinder filled with approximately 3 Torr of xenon and 4 Torr of neon. The ends of the cylinder are angled approximately 10 degrees from a perpendicular to the electrode axis; though this is less than Brewster's angle for a glass-air interface, it reduces the likelihood of retroreflected beams entering the discharge. This design permits unobstructed passage of a laser beam along the axis shared by the cylinder and electrodes.

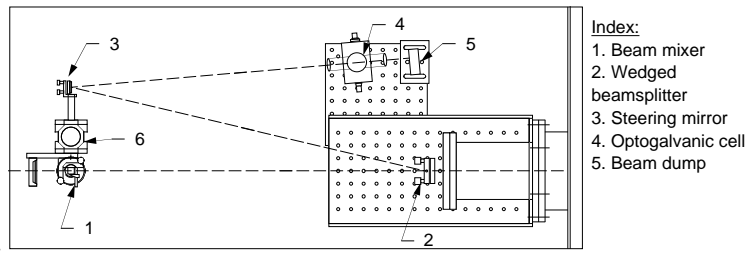


Figure 5. Optogalvanic cell setup on the CTF, plan view.

A 250 V discharge across the galvatron electrodes produces a reasonably cool, dense, stationary plasma. A ballast resistor prevents runaway current growth after the discharge is struck, and provides a simple method of measuring the discharge current. The voltage drop across the ballast resistor is connected to the Stanford SR-810 lock-in amplifier through a $3.33 \mu\text{F}$, 450 V capacitor. The capacitor passes the AC component of the voltage drop, while isolating the lock-in from high DC voltages.

D. Retarding potential analyzer

The RPA used in this experiment is based on the three-grid AFRL design described by Beal.²⁶ The outer body of the RPA is a grounded $\phi 38 \times 32$ mm shell of 1 mm thick 316 stainless steel. A phenolic sleeve isolates the grids, which are identical $\phi 32 \times 0.125$ mm disks of 316 stainless steel sheet with an etched hexagonal pattern of $\phi 0.279$ mm apertures. The total open area fraction for each grid is 38%. Macor washers fix the grid separation distances. In a departure from the original AFRL design, the collector is a $\phi 32 \times 1$ mm disk of 316 stainless steel, plasma-sprayed with a tungsten coating to minimize secondary electron emission.

During operation, the outermost grid floats, minimizing perturbations at the probe-plasma interface. A laboratory power supply holds the second (or electron repulsion) grid 30 volts below ground. A Keithley 2410 Sourcemeter (under LabView control through a GPIB interface) then sweeps the third (or ion repulsion) grid potential from 0 to 600 V. A Keithley 6485 picoammeter simultaneously measures and passes the collector current via GPIB to the LabView VI. Numerical differentiation of the collected current as a function of ion retarding grid potential provides an instrument trace that is linearly proportional to the distribution of ion energy per unit charge. For the sake of brevity, we will henceforth refer to this as an ion voltage distribution function.²⁶

IV. Results

A. OGS at 605.1 nm

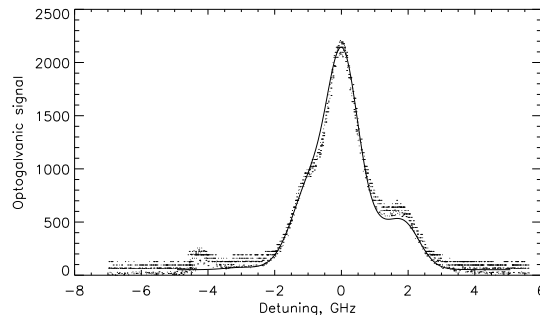


Figure 6. Optogalvanic spectrum at 605.1 nm (points) and warm-plasma spectrum model, $T = 1255$ K (line).

Figure 6 shows a sample set of seven optogalvanic spectra taken at 605.1 nm with approximately 8%

of the maximum dye laser output. From Maxwellian curve-fitting to velocity distributions deconvolved from 49 such data sets, we estimate that the Xe II temperature in a 250 V optogalvanic cell discharge is $T = (1255 \pm 57)$ K. The solid line in Figure 6 shows a scaled warm-plasma spectrum model at $T = 1255$ K, using the published HFS constants and the natural upper-state lifetime (7.8 ns) for the 605.1 nm absorption line^a.

The fit between the model and the data set is generally quite good. The minor peaks in the model tend to fall slightly under the experimental points. This may be evidence of saturation, which tends to preferentially flatten the highest peaks of an absorption line.

The observed signal-to-background ratio for this data set is about 32. No special efforts were made to minimize the background level during this series of experiments.

B. OGS at 834.7 nm

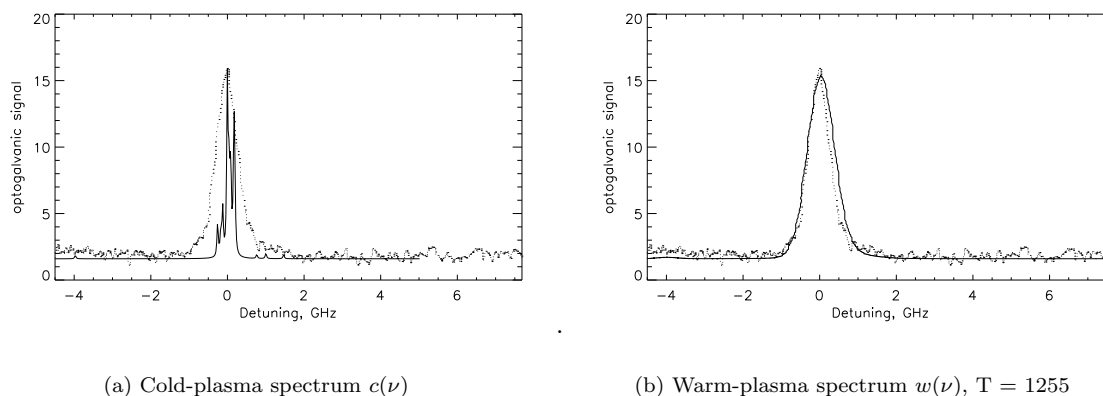


Figure 7. Optogalvanic spectrum at 834.7 nm: data (points) and spectrum model (line).

Figure 7 shows the ensemble average of 64 optogalvanic spectra taken at 834.7 nm with a 110 μ W diode laser beam at the galvatron entrance. We took especial pains to minimize the photoelectric effect for this data set, yielding a modest signal-to-background ratio of about 8. Beam quality (*i.e.* the level of departure from a perfect Gaussian beam profile), which is generally far worse for diode lasers than dye lasers, may be one culprit; however, placing a field stop in front of the galvatron did not yield impressive improvements in the signal-to-background level. The relative difficulty in achieving a good signal-to-background ratio at 834.7 nm, when compared to the ease with which we achieved far better results at 605.1 nm, suggests that the optogalvanic effect at 834.7 nm is relatively weak. Since both lines exhibit strong emission in discharge plasmas, the weak optogalvanic effect implies that the difference in upper-state and lower-state ionization probabilities is far lower for the 834.7 nm line than the 605.1 nm line.

This ensemble average served as the experimental data set for conjugate-gradient optimization of the computational model. Figure 7(a) overlays the data set with the best-fit cold-plasma spectrum, using the published natural upper-state (9.3 ns) to set the Lorentzian linewidth. Figure 7(b) overlays the same data set with a best-fit warm-plasma spectrum, using the ion temperature calculated from the 605.1 nm data to set the Doppler broadening.

The fit between the model and the data set is not as good as the fit in Figure 6. The model is uniformly broader than the data set. One possible explanation for this excessive broadening is an error in the broadening model. Our knowledge of the linewidths for broadening mechanisms can be impugned; not only is the galvatron pressure far higher than typically acceptable for purely natural broadening, but the saturation effects noted in the 605.1 nm optogalvanic spectra may exaggerate the reference cell temperature. Unfortunately, we currently have insufficient information to determine the validity of either hypothesis; further validation from the existing data set must rest on how well ion energy distributions deconvolved from 834.7 nm LIF data compare with RPA data.

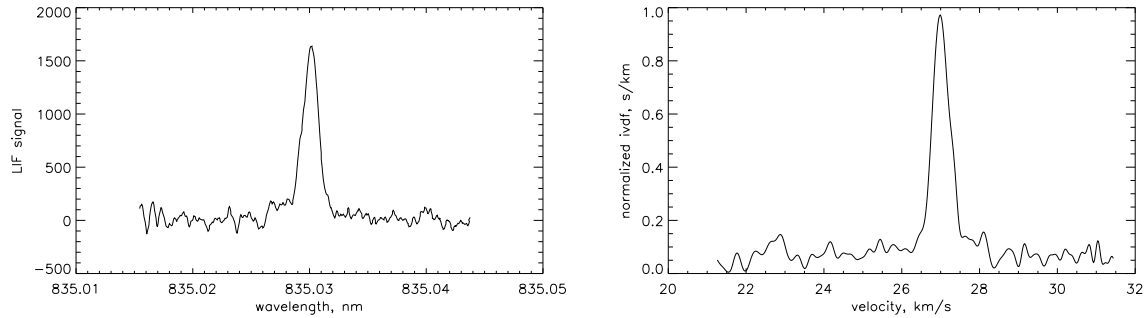
^aGiven the galvatron cold-gas pressure of 7 Torr, the Lorentzian broadening is probably dominated by collisional, rather than radiative, effects. In the absence of better information, we have opted to use the natural linewidth for this study.

Table 3 compares the previously-described “first guess” \vec{x}_0 with the final set of isotopic shifts and lower-state ($5d\ ^2F_{7/2}$) nuclear-spin constants returned by conjugate-gradient optimization. The similarity between the two sets of constants is striking; with the current data set, the simple expedient of choosing lower-state nuclear-spin constants to minimize the hyperfine-broadened linewidth comes very near to an optimal estimate. The error function for the first-guess set of constants is only 0.87% worse than the final estimate! Subsequent reconvergence studies in which we varied the Xe II temperature by the standard deviation (55 K) of the 605.1 nm temperatures resulted in isotopic shift and nuclear-spin constant uncertainties on the order of 15 MHz.

Table 3. Isotopic shifts and $5d\ ^2F_{7/2}$ nuclear-spin constants from 834.7 nm optogalvanic spectra

Solution	A_{129} (MHz)	A_{131} (MHz)	B_{131} (MHz)	$\Delta\nu_{128}$ (MHz)	$\Delta\nu_{129}$ (MHz)	$\Delta\nu_{130}$ (MHz)	$\Delta\nu_{131}$ (MHz)	$\Delta\nu_{134}$ (MHz)	$\Delta\nu_{136}$ (MHz)
First guess	-1030.5	297.2	13.2	170.73	92.379	73.461	-17.581	86.655	43.282
Convergence	-1035.4	316.3	14.33	171.25	92.97	74.257	-19.225	89.753	44.747

C. LIF velocimetry



(a) LIF spectrum.

(b) Ion velocity distribution from LIF.

Figure 8. LIF spectrum and ion velocity distribution at 834.7 nm in a 500 V, 30 mA ion beam.

Figure 8(a) shows the ensemble average of 62 LIF spectra taken at 834.7 nm in the 500 V ion source plume. Each 120-second scan consists of 501 data points taken with a lock-in time constant of 300 ms. After transmission losses, the diode laser power entering the interrogation volume dropped from 47.2 mW (at the high- λ end of the scan) to 24.7 mW (at the low- λ end of the scan). During these tests, we set an anode potential of 500 V, a discharge voltage of 32.0 V, an accel grid potential of -75.0 V, and a xenon volume flow rate of 2.0 sccm. Manual adjustment of the cathode filament current maintained the beam current at 30.0 ± 0.3 mA. Charge neutralization in the beam was ensured by increasing the neutralizer filament current until neutralization arcs were no longer observed.

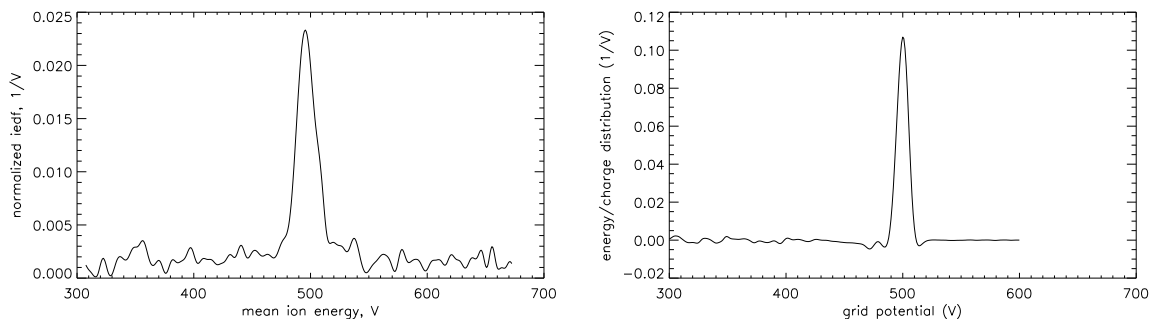
Over the course of this test, the discharge current required to maintain the beam current at 30 mA dropped unsteadily from 638 mA to 145 mA. Isolation checks showed that the accelerator grid isolation dropped from an unreadably high pre-test value to a post-test value of 50.0 k Ω . Post-test examination of the ion source showed severe damage to the accelerator grid. A new hexagonal set of $\phi 0.75$ mm apertures, offset approximately 0.75 mm from the original grid apertures, had been etched through the accelerator grid. We believe this to be caused by misalignment of the grids during reassembly, as the isolators had just been removed for cleaning before the test.

Figure 8(b) shows the ion velocity distribution deconvolved from these LIF spectra. The deconvolution routine uses the newly-modeled 834.7 nm cold-plasma spectrum, a Gaussian inverse filter with a 2.0 ns

bandwidth, and the optogalvanic spectrum peak of 834.955 nm (vacuum) as the stationary plasma reference wavelength. The mode ion speed for this distribution is 26.985 km/s, while the mean ion speed for the primary peak (defined as the area between the two minima on either side of the ivdf maximum) is $\langle v \rangle = 26.890$ km/s. The central moments of the principal peak, defined as

$$u^{(k)} = \langle (v - \langle v \rangle)^k \rangle, \quad (45)$$

are $u^{(2)} = 4.6132$ (km/s)², $u^{(3)} = -2.7830$ (km/s)³, and $u^{(4)} = 67.507$ (km/s)⁴. These moments would equal zero for a pure Maxwellian distribution; for the measured distribution, they indicate that the peak is left-skewed ($u^{(3)} < 0$) and has mild leptokurtosis ($u^{(4)} > 0$). Since $u^{(2)} = kT/M$ for a Maxwellian distribution, the equivalent temperature for the primary peak is an astonishingly large 6.28 eV.



(a) Ion energy distribution from LIF.

(b) Ion voltage distribution from RPA.

Figure 9. LIF and RPA ion energy distributions in a 500 V ion beam.

Figure 9(a) shows the same distribution converted to an ion energy basis, while Figure 9(b) shows the ion voltage distribution from previous RPA testing in a 500 V ion beam. The ion-retarding grid potential was raised in 1-V steps during this scan, while the electron-retarding grid potential was held at a constant 30 V below ground. During the RPA test, we set an anode potential of 500 V, a discharge voltage of 32.0 V, an accel grid potential of -75.0 V, and a xenon volume flow rate of 3.0 sccm. Manual adjustment of the cathode filament current maintained the beam current at 30.0 ± 0.3 mA. Charge neutralization in the beam was ensured by increasing the neutralizer filament current until neutralization arcs were no longer observed. Over the course of the test, the discharge current required to maintain this beam current remained steady.

Table 4. Central moments of the ion energy distribution function main peak at 500 V.

	mode E	$E^{(1)}$	$E^{(2)}$	$E^{(3)}$	$E^{(4)}$
	(V)	(V)	(V ²)	(V ³)	(V ⁴)
LIF	495.45	495.10	6059.4	-12884.	1.1389×10^8
RPA	500.	499.88	18.957	-5.8123	879.82

Table 4 compares mode energies and the central moments of the principal peak for both distributions. As before, we define the k^{th} central moment of the energy distribution as

$$E^{(k)} = \langle (E - \langle E \rangle)^k \rangle. \quad (46)$$

The mode and mean energies for both measurements agree within 1%, and the third moments show that both distributions are left-skewed ($E^{(3)} < 0$). Otherwise, though, the higher moments point out that the LIF distribution is far wider and has much wider tails than the RPA distribution. As mentioned above, the equivalent temperature for an equally wide Maxwellian is 6.28 eV, which is at least an order of magnitude hotter than the axial Xe II temperatures we have previously measured in the P5 Hall thruster plume with 605.1 nm LIF.¹

We believe (but can not yet prove) that the aforementioned grid misalignment is to blame for the unnaturally wide LIF distribution. Heavily-eroded grids lead to electron backstreaming, as indicated by the decrease in discharge current over the course of the test. Both grid misalignment and electron backstreaming tend to increase the beam divergence. Williams²⁷ notes that perfectly-cold ions ($f(v) = \delta(v - \langle v \rangle)$) passing through an interrogation volume with a beam divergence angle θ will have a Doppler-broadened absorption spectrum, with an apparent temperature of

$$T_{\text{div}} = \frac{1}{2 \ln 2} \frac{M}{k} \left(\frac{\langle v \rangle}{2} \left[1 - \cos \frac{\theta}{2} \right] \right)^2. \quad (47)$$

Thus, the divergence angle required for cold ions to produce an observed velocity distribution with the variance $\langle (v - \langle v \rangle)^2 \rangle = kT_{\text{div}}/M$ is

$$\theta = 2 \cos^{-1} \left(1 - \frac{2}{\langle v \rangle} \sqrt{2 \ln 2 \langle (v - \langle v \rangle)^2 \rangle} \right). \quad (48)$$

At $\langle v \rangle = 26.9$ km/s, a beam divergence angle of $\theta = 35.7^\circ$ would produce the value of $\langle (v - \langle v \rangle)^2 \rangle$ seen in the primary peak of the LIF deconvolution.

V. Conclusions

To the best of our knowledge, this study presents the first published set of isotopic shift and lower-state nuclear-spin interaction constant estimates for the $5d \ ^2F_{7/2} \rightarrow 6p \ ^2D_{5/2}^0$ absorption at 834.7 nm. Though the optogalvanic signal-to-background ratio at 834.7 nm was significantly worse than the signal-to-background ratio for the $5d^4 D_{7/2} \rightarrow 6p^4 P_{5/2}^0$ absorption at 605.1 nm, we were able to converge on HFS constants that produce an acceptable fit to the observed optogalvanic spectra.

Our attempt to validate these HFS constants by deconvolution of 834.7 nm LIF spectra was less successful. Aside from the readily-obtained agreement between mean ion energies, the similarities between LIF and RPA ion energy distributions were more qualitative than quantitative. The expected reduction in Doppler broadening from kinematic compression effect did not appear; instead, the LIF velocity distribution was unusually wide. Misalignment of the ion source grids may be responsible for the high width of the LIF velocity distribution.

Future tests of this technique should involve simultaneous LIF and RPA measurements, as a minimum; this would permit early detection of beam divergence effects. A higher-density ion source (such as a Hall thruster) might reduce the extensive integration times required for an acceptable signal-to-noise ratio with a gridded ion source, as would replacing the neutralizer filament with a neutralizing hollow cathode. Ideally, the velocity distribution would be independently and non-intrusively measured by two-wavelength simultaneous LIF (TWSLIF) at 605.1 and 834.7 nm.

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